

arcesetc: ARC Echelle Spectrograph Exposure Time Calculator

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Software

- [Review](#) ↗
- [Repository](#) ↗
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Summary

The ARC Echelle Spectroscopic (ARCES) Exposure Time Calculator, or `arcesetc`, is a simple exposure time calculator for the ARCES instrument on the Astrophysical Research Consortium (ARC) 3.5 m Telescope at Apache Point Observatory for stellar spectroscopy. Astronomers can use it to plan observations with the ARCES instrument. Users can supply `arcesetc` functions with the spectral type of their target star, the V band magnitude, and either: the desired exposure time in order to determine the counts and signal-to-noise ratio as a function of wavelength; or the desired signal-to-noise ratio at a given wavelength to determine the required exposure time.

We estimate the count rates for stars as a function of wavelength by fitting 15th-order polynomials to each spectral order of real observations of a star of each spectral type. These polynomial coefficients and some wavelength metadata are stored in an HDF5 archive for compactness and ease of reconstruction. Then upon calling `arcesetc`, the archive is opened and the spectral order closest to the wavelength of interest is reconstructed from the polynomial coefficients, for a star of the closest available spectral type to the one requested.

At present, the 79 stellar spectral types included in the `arcesetc` library span from mid F to mid M stars on the main sequence, a variety of M giants, a handful of O and B, and a white dwarf and a Wolf-Rayet star. Contributions from the community are welcome to expand the library to include other spectral types.

`arcesetc` was built from the Astropy package-template, and thus includes self-building documentation and continuous integration (Astropy Collaboration et al., 2018).

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References

Astropy Collaboration, Price-Whelan, A. M., Sipőcz, B. M., Günther, H. M., Lim, P. L., Crawford, S. M., Conseil, S., et al. (2018). The Astropy Project: Building an Open-science Project and Status of the v2.0 Core Package, *156*, 123. doi:[10.3847/1538-3881/aabc4f](https://doi.org/10.3847/1538-3881/aabc4f)