aiapy: A Python Package for Analyzing Solar EUV Image Data from AIA

Will T. Barnes¹, ², Mark C. M. Cheung², Monica G. Bobra³, Paul F. Boerner², Georgios Chintzoglou², ⁴, Drew Leonard⁵, Stuart J. Mumford⁶, Nicholas Padmanabhan², ⁷, Albert Y. Shih⁸, Nina Shirman², David Stansby⁹, and Paul J. Wright⁴

¹ National Research Council Postdoctoral Research Associate residing at the Naval Research Laboratory, Washington, D.C. 20375, USA ² Lockheed Martin Solar and Astrophysics Laboratory, Palo Alto, CA 94304, USA ³ W. W. Hansen Experimental Physics Laboratory, Stanford University, Stanford, CA 94305, USA ⁴ University Corporation for Atmospheric Research, Boulder, CO 80301, USA ⁵ Aperio Software Ltd, Leeds LS6 3HN, UK ⁶ School of Mathematics and Statistics, The University of Sheffield, Sheffield S3 7RH, UK ⁷ Princeton University, Princeton, NJ 08544, USA ⁸ NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA ⁹ Mullard Space Science Laboratory, University College London, Holmbury St. Mary, Surrey RH5 6NT, UK

DOI: 10.21105/joss.02801

Summary

The Atmospheric Imaging Assembly (AIA; Lemen et al., 2012) instrument onboard the NASA Solar Dynamics Observatory (SDO; Pesnell et al., 2012) spacecraft has observed the full-disk of the Sun, nearly continuously, for the last ten years. It is one of three instruments on SDO, along with the Helioseismic and Magnetic Imager (HMI; Schou et al., 2012) and the Extreme Ultraviolet Variability Experiment (EVE; Woods et al., 2012). With its high spatial (0.6” per pixel) and temporal (up to 12 seconds for most channels) resolution, AIA has greatly enhanced our understanding of our closest star in a number of different areas, including the initiation of flares and coronal mass ejections as well as the quiescent heating of the corona, the outermost layer of the Sun’s atmosphere. AIA is a narrowband imaging instrument comprised of four separate telescopes that collectively observe the full-disk of the Sun at ten different wavelengths: seven extreme ultraviolet (EUV) wavelengths, two far UV wavelengths, and one visible wavelength. It produces nearly 60,000 images per day with a new 4K-by-4K image produced by each EUV channel every 12 seconds. The image data are provided to the community by the Joint Science Operations Center (JSOC) at Stanford University in the Flexible Image Transport System (FITS; Wells et al., 1981) format.

Statement of Need

aiapy is a Python package for analyzing calibrated (level 1) EUV imaging data from AIA. It includes capabilities for aligning images between channels, deconvolving images with the instrument point-spread function (PSF), computing channel sensitivity as a function of wavelength, and correcting images for telescope degradation, among others. Historically, most data analysis in solar physics has been done using the Interactive Data Language (IDL), a propriety interpreted language commonly used throughout astronomy. The AIA instrument team has developed and continues to maintain a comprehensive set of IDL tools for querying,

*Other than the first two authors, the author list in this paper is sorted alphabetically.*
calibrating, and analyzing AIA data that has been widely used for nearly a decade. A full description of this software can be found in the SDO Data Analysis Guide.

As the solar physics community, and the astronomy community as a whole, began transitioning to Python, some of this functionality was absorbed into sunpy, the core Python package for solar data analysis (SunPy Project et al., 2020). As the sunpy package grew and Python adoption amongst solar physicists increased (see Bobra et al., 2020), the need for instrument-specific packages maintained by instrument teams has become apparent. As such, the aim of aiapy is to provide instrument-specific functionality while being fully interoperable with both the SunPy and broader Astropy (The Astropy Collaboration et al., 2018) ecosystems.

**Figure 1**: Some examples of the capabilities of the aiapy package. The top row shows a cutout of an AIA 171 Å image before (top left) and after (top right) being convolved with the instrument point spread function (PSF). The bottom left panel shows the degradation as a function of time for all seven EUV channels since the launch of SDO. The bottom right panel shows the wavelength response functions for all seven EUV channels.

### Package Structure

aiapy has three primary subpackages. Figure 1 shows several examples of the functionality included in these subpackages. The aiapy.calibrate subpackage contains functions for correcting images for telescope degradation, normalizing to the exposure time, and replacing
“hot” pixels removed during earlier calibration procedures. Additionally, this subpackage also contains the register function for removing the roll angle, aligning the center of the image with the center of the Sun, and scaling the image to a common resolution across channels. An image that has been rotated, aligned, and rescaled in this manner is a level 1.5 image. This function, combined with the ability to update the image metadata that describes the satellite pointing provided by the update_pointing function, replicates the commonly-used aiaprep.pro IDL procedure, provided in the SolarSoftware ecosystem (Freeland & Handy, 1998), used to align images across channels.

The aiapy.psf subpackage includes functions for calculating the instrument point spread function (PSF) for each EUV channel and deconvolving images with the PSF via Richardson-Lucy deconvolution. Because computing the PSF and performing the deconvolution is computationally expensive for a full-frame AIA image, aiapy.psf also includes optional GPU acceleration via cupy (Okuta et al., 2017). The top row of Figure 1 shows an example of a 171 Å image that has been deconvolved with the PSF.

Finally, the aiapy.response subpackage provides the Channel class for computing the wavelength response functions for each channel as a function of wavelength. Optional corrections to the response functions for instrument degradation and channel crosstalk are also provided. Additionally, individual components of the wavelength response, including the primary and secondary reflectance of each telescope as well as the efficiency of the secondary and focal plane filters, are accessible via the Channel class. The AIA wavelength response functions are described in detail in Boerner et al. (2012). Combined with atomic data from the CHIANTI atomic database (Dere et al., 1997), these wavelength response functions can be used to compute the temperature sensitivity of each EUV channel. The lower right panel of Figure 1 shows the seven EUV wavelength response functions as a function of wavelength.

Development and Infrastructure

Version 0.3.0 of aiapy was released on 6 October 2020 and is available through the Python Package Index via pip. aiapy is compatible with Python 3.6+ and is built on top of sunpy and astropy and utilizes the drms package (Glogowski et al., 2019) for retrieving metadata information from the JSOC. aiapy is also part of the SunPy ecosystem and is a SunPy-affiliated package (See Section 6 of SunPy Project et al., 2020). As such, it is reviewed on an annual basis to ensure it meets certain established criteria in the categories of functionality, documentation, testing, and community engagement, among others. Additionally, the code is developed openly on GitLab and the documentation is hosted online on Read the Docs. aiapy includes a comprehensive test suite built on top of thepytest testing framework. The full test suite, including tests for all supported versions of Python, online tests, documentation builds, and code style checks, is run on every single code contribution using the built-in continuous integration pipelines in GitLab and test coverage is monitored and reported using Codecov. This test suite is also run weekly to monitor any failures that may occur due to upstream changes in package dependencies.

Acknowledgements

The authors acknowledge support from NASA’s SDO/AIA contract (NNG04EA00C) to the Lockheed Martin Solar and Astrophysics Laboratory. AIA is an instrument onboard the Solar Dynamics Observatory, a mission for NASA’s Living With a Star program. WTB was supported by NASA’s Hinode program. Hinode is a Japanese mission developed and launched by ISAS/JAXA with NAOJ as a domestic partner and NASA and STFC (UK) as international partners. It is operated by these agencies in cooperation with ESA and NSC (Norway).
GC acknowledges support by NASA HSR grant 80NSSC19K0855. DS is supported by STFC grant ST/S000240/1. PJW acknowledges support from NASA Contract NASS-02139 (HMI) to Stanford University.

References


